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# **Hidden Symmetries of Charged Kerr Black Hole**

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**Tsuyoshi Houri ( Osaka City University )**

\* collaboration with

D.Kubizňák, C.M.Warnick (DAMTP) and Y.Yasui (OCU)

\* in preparation

## Motivation

- String theory implies the existence of extra dimensions and motivates us to study a gravity in a higher-dimensional framework.
- There is gravity/gauge duality which is one of the most exciting ideas in particle physics.

$$(d + 1)\text{-dim. gravitaional theory} \Leftrightarrow d\text{-dim. gauge theory}$$

- Understanding in higher-dimensional framework might give us further understanding in 4-dimension.

Black hole solutions provide important and useful gravitational back-grounds for these purposes, since black holes possess properties such as entropy and a singularity that fundamental physics aims to address.

## Black hole metrics in a vacuum

- 4-dimensional black hole metric

	mass	a.m.	NUT	$\Lambda$
Schwarzschild (1915)				
Kerr (1963)				
Carter (1968)				
Plebanski (1975)				

- Higher-dimensional ( $D \geq 4$ ) black hole metric

	mass	a.m.s	NUTs	$\Lambda$
Tangherlini (1963)				
Myers, Perry (1986)				
Gibbons, Lü, Page, Pope (2004)				
<b>Chen, Lü, Pope (2006)</b>				

## Kerr-NUT-AdS metric in $D$ -dimension

The most general known solution (Chen-Lü-Pope metric) is called **Kerr-NUT-AdS metric**, which is given by

$$g = \sum_{\mu=1}^n \frac{dx_{\mu}^2}{Q_{\mu}} + \sum_{\mu=1}^n Q_{\mu} \left[ \sum_{k=0}^{n-1} A_{\mu}^{(k)} d\psi_k \right]^2 + \varepsilon S \left[ \sum_{k=0}^n A^{(k)} d\psi_k \right]^2$$

in  $D = 2n + \varepsilon$  dimension, where  $\varepsilon = 0$  for even dimensions and  $\varepsilon = 1$  for odd dimensions.

Here the functions are

$$Q_{\mu} = \frac{X_{\mu}}{U_{\mu}}, \quad U_{\mu} = \prod_{\nu \neq \mu} (x_{\mu}^2 - x_{\nu}^2), \quad X_{\mu} = \sum_{k=\varepsilon}^n c_k x_{\mu}^{2k} + b_{\mu} x_{\mu}^{1-\varepsilon} + \varepsilon \frac{(-1)^k c}{x_{\mu}^2},$$

$$A_{\mu}^{(k)} = \sum_{\substack{1 \leq \nu_1 < \dots < \nu_k \leq n \\ \nu_i \neq \mu}} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A^{(k)} = \sum_{1 \leq \nu_1 < \dots < \nu_k \leq n} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A_{\mu}^{(0)} = A^{(0)} = 1,$$

$$S = \frac{c}{A^{(n)}}, \quad c = \text{const.}$$

This metric satisfies  $R_{ab} = -(D - 1)c_n g_{ab}$  in all dimensions.

Kerr metric (4-dimension)

$$ds^2 = \frac{\Sigma}{\Delta} dr^2 + \Sigma d\theta^2 - \left( \frac{\Delta - a^2 \sin^2 \theta}{\Sigma} \right) dt^2 - \frac{4Mar \sin^2 \theta}{\Sigma} dt d\phi + \left[ \frac{(r^2 + a^2)^2 - \Delta a^2 \sin^2 \theta}{\Sigma} \right] \sin^2 \theta d\phi^2$$

where

$$\Sigma = r^2 + a^2 \sin^2 \theta, \quad \Delta = r^2 + a^2 - 2Mr$$

Kerr metric (4-dimension)

$$ds^2 = \frac{x^2 - y^2}{X} dx^2 + \frac{y^2 - x^2}{Y} dy^2 \\ + \frac{X}{x^2 - y^2} (d\psi_0 + y^2 d\psi_1)^2 + \frac{Y}{y^2 - x^2} (d\psi_0 + x^2 d\psi_1)^2$$

where

$$X = x^2 - a^2 - 2Mx, \quad Y = y^2 - a^2$$

.....  
Kerr-NUT metric (4-dimension)

$$ds^2 = \frac{x^2 - y^2}{X} dx^2 + \frac{y^2 - x^2}{Y} dy^2 \\ + \frac{X}{x^2 - y^2} (d\psi_0 + y^2 d\psi_1)^2 + \frac{Y}{y^2 - x^2} (d\psi_0 + x^2 d\psi_1)^2$$

where

$$X = x^2 - a^2 - 2Mx, \quad Y = y^2 - a^2 - 2Ly$$

Ansatz metric (4-dimension)

$$ds^2 = \frac{x^2 - y^2}{X(x)} dx^2 + \frac{y^2 - x^2}{Y(y)} dy^2 \\ + \frac{X(x)}{x^2 - y^2} (d\psi_0 + y^2 d\psi_1)^2 + \frac{Y(y)}{y^2 - x^2} (d\psi_0 + x^2 d\psi_1)^2$$

We can determine the functions  $X$  and  $Y$  by imposing Einstein condition  $R_{ab} = -3c g_{ab}$ .

.....

Kerr-NUT-AdS metric (4-dimension)

$$ds^2 = \frac{x^2 - y^2}{X} dx^2 + \frac{y^2 - x^2}{Y} dy^2 \\ + \frac{X}{x^2 - y^2} (d\psi_0 + y^2 d\psi_1)^2 + \frac{Y}{y^2 - x^2} (d\psi_0 + x^2 d\psi_1)^2$$

where

$$X = cx^4 + x^2 - a^2 - 2Mx, \quad Y = cy^4 + y^2 - a^2 - 2Ly$$

Kerr-NUT-AdS metric (5-dimension)

$$\begin{aligned} ds^2 = & \frac{x^2 - y^2}{X} dx^2 + \frac{y^2 - x^2}{Y} dy^2 \\ & + \frac{X}{x^2 - y^2} (d\psi_0 + y^2 d\psi_1)^2 + \frac{Y}{y^2 - x^2} (d\psi_0 + x^2 d\psi_1)^2 \\ & + \frac{c}{x^2 y^2} (d\psi_0 + (x^2 + y^2) d\psi_1 + x^2 y^2 d\psi_2)^2 \end{aligned}$$

$$X = c_4 x^4 + c_2 x^2 + c_0 + b_1 + \frac{c}{x^2}, \quad Y = c_4 y^4 + c_2 y^2 + c_0 + b_2 + \frac{c}{y^2}$$



Kerr-NUT-AdS metric (6-dimension)

$$\begin{aligned}
 ds^2 = & \frac{(x^2 - y^2)(x^2 - z^2)}{X} dx^2 + \frac{(y^2 - x^2)(y^2 - z^2)}{Y} dy^2 + \frac{(z^2 - x^2)(z^2 - y^2)}{Z} dz^2 \\
 & + \frac{X}{(x^2 - y^2)(x^2 - z^2)} (d\psi_0 + (y^2 + z^2)d\psi_1 + y^2 z^2 d\psi_2)^2 \\
 & + \frac{Y}{(y^2 - x^2)(y^2 - z^2)} (d\psi_0 + (z^2 + x^2)d\psi_1 + z^2 x^2 d\psi_2)^2 \\
 & + \frac{Z}{(z^2 - x^2)(z^2 - y^2)} (d\psi_0 + (x^2 + y^2)d\psi_1 + x^2 y^2 d\psi_2)^2
 \end{aligned}$$

where

$$\begin{aligned}
 X &= c_6 x^6 + c_4 x^4 + c_2 x^2 + c_0 + b_1 x , \\
 Y &= c_6 y^6 + c_4 y^4 + c_2 y^2 + c_0 + b_2 y , \\
 Z &= c_6 z^6 + c_4 z^4 + c_2 z^2 + c_0 + b_3 z
 \end{aligned}$$

## Kerr-NUT-AdS metric (7-dimension)

$$\begin{aligned}
 ds^2 = & \frac{(x^2 - y^2)(x^2 - z^2)}{X} dx^2 + \frac{(y^2 - x^2)(y^2 - z^2)}{Y} dy^2 + \frac{(z^2 - x^2)(z^2 - y^2)}{Z} dz^2 \\
 & + \frac{X}{(x^2 - y^2)(x^2 - z^2)} (d\psi_0 + (y^2 + z^2)d\psi_1 + y^2 z^2 d\psi_2)^2 \\
 & + \frac{Y}{(y^2 - x^2)(y^2 - z^2)} (d\psi_0 + (z^2 + x^2)d\psi_1 + z^2 x^2 d\psi_2)^2 \\
 & + \frac{Z}{(z^2 - x^2)(z^2 - y^2)} (d\psi_0 + (x^2 + y^2)d\psi_1 + x^2 y^2 d\psi_2)^2 \\
 & + \frac{c}{x^2 y^2 z^2} (d\psi_0 + (x^2 + y^2 + z^2)d\psi_1 + (x^2 y^2 + y^2 z^2 + x^2 z^2)d\psi_2 + x^2 y^2 z^2 d\psi_3)^2
 \end{aligned}$$

where

$$\begin{aligned}
 X &= c_6 x^6 + c_4 x^4 + c_2 x^2 + c_0 + b_1 - \frac{c}{x^2} , \\
 Y &= c_6 y^6 + c_4 y^4 + c_2 y^2 + c_0 + b_2 - \frac{c}{y^2} , \\
 Z &= c_6 z^6 + c_4 z^4 + c_2 z^2 + c_0 + b_3 - \frac{c}{z^2}
 \end{aligned}$$

We can assume the ansatz metric

$$g = \sum_{\mu=1}^n \frac{dx_{\mu}^2}{Q_{\mu}} + \sum_{\mu=1}^n Q_{\mu} \left[ \sum_{k=0}^{n-1} A_{\mu}^{(k)} d\psi_k \right]^2 + \varepsilon S \left[ \sum_{k=0}^n A^{(k)} d\psi_k \right]^2$$

in  $D = 2n + \varepsilon$  dimension, where  $\varepsilon = 0$  for even dimensions and  $\varepsilon = 1$  for odd dimensions.

Here the functions are

$$Q_{\mu} = \frac{X_{\mu}}{U_{\mu}}, \quad U_{\mu} = \prod_{\nu \neq \mu} (x_{\mu}^2 - x_{\nu}^2), \quad X_{\mu} = X_{\mu}(x_{\mu}),$$

$$A_{\mu}^{(k)} = \sum_{\substack{1 \leq \nu_1 < \dots < \nu_k \leq n \\ \nu_i \neq \mu}} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A^{(k)} = \sum_{1 \leq \nu_1 < \dots < \nu_k \leq n} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A_{\mu}^{(0)} = A^{(0)} = 1,$$

$$S = \frac{c}{A^{(n)}}, \quad c = \text{const.}$$

Imposing Einstein condition  $R_{ab} = \lambda g_{ab}$ , we can determine the form of the function  $X_{\mu}$

$$X_{\mu} = \sum_{k=\varepsilon}^n c_k x_{\mu}^{2k} + b_{\mu} x_{\mu}^{1-\varepsilon} + \varepsilon \frac{(-1)^k c}{x_{\mu}^2}.$$

## Separabilities of Kerr-NUT-AdS spacetime in higher-dimensions

It is known that the separation of variables for various field equations on Kerr-NUT-AdS background.

- Geodesic equation

Frolov-Krtous-Kubiznak-Page(2006)

- Klein-Gordon equation

Kubiznak-Krtous-Kubiznak(2006)

- Dirac equation

Oota-Yasui(2008), Wu(2009)

- gravitational perturbation equation (tensor modes)

Kundri-Lucietti-Reall(2006), Oota-Yasui(2008)

- Maxwell equation ?

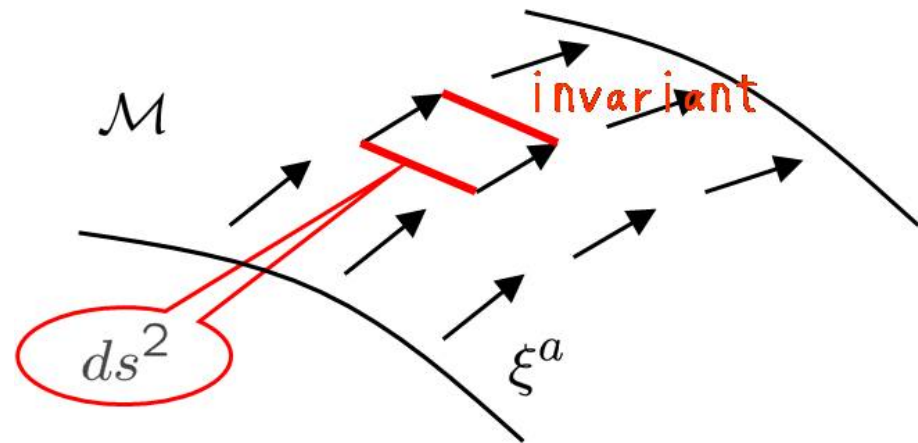
## Killing vector

Def. A generator of isometry of spacetime  $\xi$ , i.e.,

$$\nabla_{(a}\xi_{b)} = 0 \quad (\mathcal{L}_\xi g = 0) ,$$

is called **Killing vector**.

If the orbit of Killing vector is closed, it generates axial symmetry. If not, it generates translation symmetry.



## Conformal Killing vector

Def. A generator of conformal symmetry of spacetime  $\xi$ , i.e.,

$$\nabla_{(a}\xi_{b)} = \phi g_{ab} \quad (\mathcal{L}_\xi g = 2\phi g) ,$$

is called **conformal Killing vector**.

## Geodesic integrability

For geodesic Hamiltonian  $H = \frac{1}{2}g_{ab}p^ap^b$ , E.O.M. gives geodesic equation

$$p^b\nabla_b p^a = 0 \quad (\ddot{x}^a + \Gamma^a_{bc}\dot{x}^b\dot{x}^c = 0) .$$

We assume that a C.O.M. is written as  $C = K_{a_1\dots a_n}p^{a_1}\dots p^{a_n}$ . Then the condition

$$\{C, H\}_P = 0$$

leads to the equation

$$\nabla_{(b}K_{a_1\dots a_n)} = 0 .$$

This equation is called **Killing equation** and  $K$  is called **Killing tensor of rank-n**. When  $n = 1$ ,  $K$  is a Killing vector.

Since Killing tensor gives C.O.M. along geodesic, geodesic equation is integrable if there are the dimension number of Killing vectors and Killing tensors totally.

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motivation

solution admitting a closed conformal Killing-Yano tensor

solution admitting a generalized closed conformal Killing-Yano tensor

summary and discussion

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symmetric	Killing tensor	conformal Killing tensor
anti-symmetric	Killing-Yano tensor	conformal Killing-Yano tensor

## Geodesic integrability of Kerr spacetime in 4-dimension

**Carter (1968)** ... There exists a nontrivial Killing tensor  $K$ , so there are four constants of motion.

$$\xi = \partial_t, \quad \eta = \partial_\phi, \quad g, \quad K$$

**Penrose and Floyd (1973)** ... Killing tensor  $K$  is written as the square of rank-2 Killing-Yano tensor  $f$ .

$$\exists f \quad \text{s.t.} \quad K_{ab} = f^c{}_a f_{bc}, \quad f_{ba} = -f_{ab}, \quad \nabla_{(a} f_{b)c} = 0$$

KY equation

**Hughston and Sommers (1987)** ... Two Killing vectors,  $\xi$  and  $\eta$ , are also constructed from the Killing-Yano tensor  $f$ .

$$\xi^a = \nabla_b (*f)^{ba}, \quad \eta^a = K^a{}_b \xi^b$$

**$\Rightarrow$  KY tensor is more fundamental.**



## Killing tensor

Def. When a rank- $n$  symmetric tensor  $K$  satisfies the equation

$$\nabla_{(b} K_{a_1 \dots a_n)} = 0 ,$$

$K$  is called **Killing tensor**.

## Killing-Yano tensor

Def. When a rank- $n$  anti-symmetric tensor  $f$  satisfies the equation

$$\nabla_{(b} f_{a_1) a_2 \dots a_n} = 0 ,$$

$f$  is called **Killing-Yano (KY) tensor**.

## Geodesic integrability of Kerr-NUT-AdS spacetime in $D$ -dimension

Page, Frolov, Kubizňák, Krtous and Vasdevan (2006)

There exist  $n - 1$  nontrivial Killing tensors  $K^{(j)}$  in  $D$ -dimension, so there are the dimension number of constants of motion, which are mutually commuting.

$$\xi = \partial_t, \quad \eta^{(j)} = \partial_{\phi_i}, \quad g, \quad K^{(j)} \quad \text{and} \quad \eta^{(n)} \quad (j = 1, \dots, n - 1)$$

# Dimension	# Killing vector	# Killing tensor
$D = 2n$	$n$	$n$
$D = 2n + 1$	$n + 1$	$n$

As the 4-dimension, Killing vectors and tensors,  $\xi$ ,  $\eta^{(j)}$  and  $K^{(j)}$ , are constructed from rank- $(D - 2j)$  Killing-Yano tensors  $f^{(j)}$ .

$$K_{ab}^{(j)} = f^{(j)}_a \dots f^{(j)}_b \dots, \quad \xi^a = \nabla_b (*f^{(1)})^{ba}, \quad \eta^{(j)a} = K^{(j)a}{}_b \xi^b$$

## Geodesic integrability of Kerr-NUT-AdS spacetime in $D$ -dimension

Futhermore,  $n - 1$  Killing-Yano tensors  $f^{(j)}$  are constructed from a single rank-2 CKY tensor  $h$ .

$$f^{(j)} = *h^{(j)} , \quad h^{(j)} = h \wedge h \wedge \dots \wedge h$$

( $j$  times)

**$\Rightarrow$  CKY tensor is the most fundamental.**

## Conformal Killing-Yano tensor

**Def.** For a rank- $n$  anti-symmetric tensor  $h$ , when there exists a rank- $(n - 1)$  anti-symmetric tensor  $\xi$  such that

$$\nabla_{(a} h_{b)c_1 \dots c_{n-1}} = g_{ab} \xi_{c_1 \dots c_{n-1}} + \sum_{i=1}^{n-1} (-1)^i g_{c_i(a} \xi_{b)c_1 \dots \hat{c}_i \dots c_{n-1}} ,$$

$h$  is called **conformal Killing-Yano (CKY) tensor** and  $\xi$  is called associated tensor of  $h$ ,

$$\xi_{c_1 \dots c_{n-1}} = \frac{1}{D - n + 1} \nabla^a h_{ac_1 \dots c_{n-1}} .$$

In particular, if  $\xi = 0$  then  $h$  is called **Killing-Yano (KY) tensor**.

Tachibana and Kashiwada (1968)

## Closed conformal Killing-Yano tensor

**Def.** Let  $h$  be a  $p$ -form. If  $h$  satisfies the equations

$$\nabla_X h = -\frac{1}{D-p+1} X^b \wedge \delta h \quad \text{and} \quad dh = 0$$

for  $\forall X \in TM$ , then we call  $h$  **rank- $p$  closed conformal Killing-Yano (CCKY) tensor**.

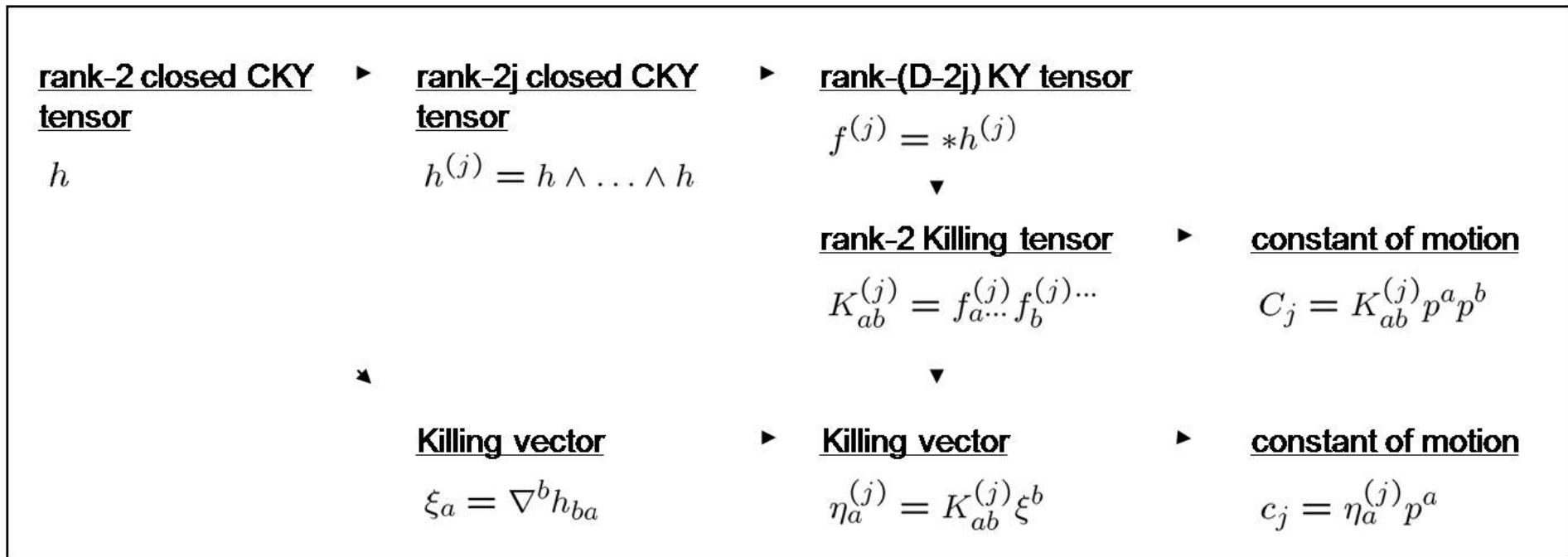
$\nabla$  : Levi-Civita connection,  $\wedge$  : wedge product,  $d$  : exterior derivative,  
 $\delta$  : coderivative operator ( $= *d*$ )

	Killing vector	conformal Killing vector
symmetric	Killing tensor	conformal Killing tensor
anti-symmetric	Killing-Yano tensor	conformal Killing-Yano tensor

**Prop.** Suppose that a spacetime admits a rank-2 non-degenerate CCKY tensor. Then the geodesic equation is integrable, namely there are the dimension number of Killing vectors and rank-2 Killing tensors totally.

Houri, Oota and Yasui (2007), Krtous, Frolov and Kubizňák (2008)

# Dimension	# Killing vector	# Killing tensor
$2n$	$n$	$n$
$2n + 1$	$n + 1$	$n$



We can prove that

$$\{C_i, C_j\}_P = 0, \quad \{C_i, c_j\}_P = 0, \quad \{c_i, c_j\}_P = 0.$$



**Theor.** We assume that a spacetime admits a rank-2 non-degenerate CCKY tensor. Then such a spacetime is given only by the metric of Kerr-NUT-AdS type. (Einstein equation is not imposed.)

Houri, Oota and Yasui (2007), Krtous, Frolov and Kubizňák (2008)

## Kerr-NUT-AdS-type metric in $D = 2n + \varepsilon$ dimension

$$g = \sum_{\mu=1}^n \frac{dx_{\mu}^2}{Q_{\mu}} + \sum_{\mu=1}^n Q_{\mu} \left[ \sum_{k=0}^{n-1} A_{\mu}^{(k)} d\psi_k \right]^2 + \varepsilon S \left[ \sum_{k=0}^n A^{(k)} d\psi_k \right]^2$$

where

$$Q_{\mu} = \frac{X_{\mu}}{U_{\mu}}, \quad U_{\mu} = \prod_{\nu \neq \mu} (x_{\mu}^2 - x_{\nu}^2), \quad X_{\mu} = X_{\mu}(x_{\mu}),$$

$$A_{\mu}^{(k)} = \sum_{\substack{1 \leq \nu_1 < \dots < \nu_k \leq n \\ \nu_i \neq \mu}} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A^{(k)} = \sum_{1 \leq \nu_1 < \dots < \nu_k \leq n} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A_{\mu}^{(0)} = A^{(0)} = 1,$$

$$S = \frac{c}{A^{(n)}}, \quad c = \text{const.}$$

# Solutions admitting a rank-2 closed CKY tensor

4-dimensional black hole metric

	mass	a.m.	NUT	$\Lambda$
Schwarzschild (1915)				
Kerr (1963)				
Carter (1968)				
Plebanski (1975)				

Higher-dimensional ( $D \geq 4$ ) black hole metric

	mass	a.m.s	NUTs	$\Lambda$
Tangherlini (1963)				
Myers, Perry (1986)				
Gibbons, Lü, Page, Pope (2004)				
Chen, Lü, Pope (2006)				

4-dimensional Kerr-Newman metric

## Theorem

We assume that  $D$ -dimensional spacetime  $(M, g)$  admits a single rank-2 closed CKY tensor. Then  $(M, g)$  is the only generalized Kerr-NUT-AdS spacetime. (Here Einstein condition is not imposed.)

Houri, Oota and Yasui (2008)

rank-2 **non-degenerate** closed conformal Killing-Yano tensor  $\implies$  unique Kerr-NUT-AdS metric

rank-2 closed conformal Killing-Yano tensor  $\implies$  unique **generalized Kerr-NUT-AdS metric**

$$\begin{aligned}
 h &= \sum_{\mu=1}^n x_{\mu} e^{\mu} \wedge e^{n+\mu} + \xi_1 \sum_{\alpha_1=1}^{m_1} e^{\alpha_1} \wedge e^{m_1+\alpha_1} + \dots + \xi_N \sum_{\alpha_N=1}^{m_N} e^{\alpha_N} \wedge e^{m_N+\alpha_N} \\
 &= \sum_{\mu=1}^n x_{\mu} e^{\mu} \wedge e^{n+\mu} + \sum_{j=1}^N \left( \xi_j \sum_{\alpha_j=1}^{m_j} e^{\alpha_j} \wedge e^{m_j+\alpha_j} \right)
 \end{aligned}$$

It is convenient to write eigenvalues of a rank-2 closed CKY tensor by introducing  $Q^a_b = -h^a_c h^c_b$ .

$$V^{-1}(Q^a_b)V = \left\{ \underbrace{-x_1^2, -x_1^2, \dots, -x_n^2, -x_n^2}_{2n}, \underbrace{-\xi_1^2, \dots, -\xi_1^2}_{2m_1}, \dots, \underbrace{-\xi_N^2, \dots, -\xi_N^2}_{2m_N}, \underbrace{0, \dots, 0}_K \right\}$$

Then  $D$ -dimensional generalized Kerr-NUT-AdS metric is

$$g = \sum_{\mu=1}^n \frac{dx_\mu^2}{P_\mu} + \sum_{\mu=1}^n P_\mu \left[ \sum_{k=0}^{n-1} A_\mu^{(k)} \theta_k \right]^2 + \sum_{j=1}^N \prod_{\mu=1}^n (x_\mu^2 - \xi_j^2) g^{(j)} + \left( \prod_{\mu} x_\mu^2 \right) g^{(0)}$$

where  $g^{(0)}$  is arbitrary  $K$ -dim. metric and  $g^{(j)}$  is  $2m_j$ -dim. Kähler metric with the Kähler form  $\omega^{(j)}$ .

$$P_\mu = \frac{X_\mu(x_\mu)}{x_\mu^K \prod_{j=1}^N (x_\mu^2 - \xi_j^2)^{m_j} \prod_{\substack{\nu=1 \\ \nu \neq \mu}}^n (x_\mu^2 - x_\nu^2)}, \quad A_\mu^{(k)} = \sum_{\substack{1 \leq \nu_1 < \dots < \nu_k \leq n \\ \nu_i \neq \mu}} x_{\nu_1}^2 \dots x_{\nu_k}^2,$$

$$d\theta_k + 2 \sum_{j=1}^N (-1)^{n-k} \xi_j^{2n-2k-1} \omega^{(j)} = 0.$$

- **$D$ -dimensional generalized Kerr-NUT-AdS metric**

$$g = \sum_{\mu=1}^n \frac{dx_{\mu}^2}{P_{\mu}} + \sum_{\mu=1}^n P_{\mu} \left[ \sum_{k=0}^{n-1} A_{\mu}^{(k)} \theta_k \right]^2 + \sum_{j=1}^N \prod_{\mu=1}^n (x_{\mu}^2 - \xi_j^2) g^{(j)} + \left( \prod_{\mu} x_{\mu}^2 \right) g^{(0)}$$

where

$$P_{\mu} = \frac{X_{\mu}(x_{\mu})}{x_{\mu}^K \prod_{j=1}^N (x_{\mu}^2 - \xi_j^2)^{m_j} \prod_{\substack{\nu=1 \\ \nu \neq \mu}}^n (x_{\mu}^2 - x_{\nu}^2)}, \quad A_{\mu}^{(k)} = \sum_{\substack{1 \leq \nu_1 < \dots < \nu_k \leq n \\ \nu_i \neq \mu}} x_{\nu_1}^2 \dots x_{\nu_k}^2,$$

$$d\theta_k + 2 \sum_{j=1}^N (-1)^{n-k} \xi_j^{2n-2k-1} \omega^{(j)} = 0.$$

When  $g^{(0)}$  is  $K$ -dim. Einstein metric,  $g^{(j)}$  is  $2m_j$ -dim. Einstein-Kähler metric with the Kähler form  $\omega^{(j)}$  and

$$X_{\mu} = x_{\mu} \int dx_{\mu} \chi(x_{\mu}) x_{\mu}^{K-2} \prod_{i=1}^N (x_{\mu}^2 - \xi_i^2)^{m_i} + d_{\mu} x_{\mu}$$

where

$$\chi(x_{\mu}) = \sum_{i=0}^n \alpha_i x_{\mu}^{2i}, \quad \alpha_0 = (-1)^{n-1} \lambda^{(0)}$$

This metric satisfies Einstein equation  $R_{ab} = -(D-1)\alpha_n g_{ab}$ .

## generalized Kerr-NUT-AdS metric

Spacetime described by generalaized Kerr-NUT-(A)dS metric has a fiber bundle structure such that

base space :        direct products of  $n$  Kähler-Einstein spaces  
fiber :                Kerr-NUT-AdS spacetime

Such a structure of spacetime appears in higher dimensional black holes with equal angular momenta.

For example,  $(2m + 3)$ -dimensional Kerr-AdS black hole metric with equal angular momenta has the follwing structure:

base space :         $CP(m)$   
fiber :                3-dimensional Kerr-NUT-AdS spacetime

# Charged rotating black holes in supergravity theory

Let us consider the following (Einstein-frame) Lagrangian :

$$\mathcal{L}_D = R * 1 + \frac{1}{2} * d\varphi \wedge d\varphi - X^{-2} * F_{(2)} \wedge F_{(2)} - \frac{1}{2} X^{-4} * H_{(3)} \wedge H_{(3)} ,$$

where

$$X = e^{-\varphi/\sqrt{2(D-2)}} , \quad F_{(2)} = dA_{(1)} , \quad H_{(3)} = dB_{(2)} - A_{(1)} \wedge dA_{(1)} .$$

This is a system consisted of gravitational field  $g$ , scalar field  $\varphi$ , 1-form potential  $A_{(1)}$  and 2-form potential  $B_{(2)}$ .

\* This Lagrangian appears as a truncation of the bosonic part of various supergravity theories, for example of heterotic supergravity compactified on a torus, and also as the ungauged limit of truncations of certain gauged supergravity theories.

# Charged Kerr-NUT solution in $D = 2n + \varepsilon$ dimension

Chow (2008)

$$g_D = H^{2/(D-2)} \left\{ \sum_{\mu=1}^n \frac{dx_\mu^2}{Q_\mu} + \sum_{\mu=1}^n Q_\mu \left( \mathcal{A}_\mu - \sum_{\nu=1}^n \frac{2N_\nu s^2}{HU_\nu} \mathcal{A}_\nu \right)^2 + \varepsilon S \left( \mathcal{A} - \sum_{\nu=1}^n \frac{2N_\nu s^2}{HU_\nu} \mathcal{A}_\nu \right)^2 \right\}$$

$$X = H^{-1/(D-2)}, \quad A_{(1)} = \sum_{\mu=1}^n \frac{2N_\mu s c}{HU_\mu} \mathcal{A}_\mu, \quad B_{(2)} = d\psi_0 \wedge \left( \sum_{\nu=1}^n \frac{2N_\nu s^2}{HU_\nu} \mathcal{A}_\nu \right).$$

Here the 1-forms and the functions are

$$\mathcal{A}_\mu = \sum_{k=0}^{n-1} A_\mu^{(k)} d\psi_k, \quad \mathcal{A} = \sum_{k=0}^n A^{(k)} d\psi_k, \quad H = 1 + \sum_{\mu=1}^n \frac{2N_\mu s^2}{U_\mu}, \quad N_\mu = m_\mu x_\mu^{1-\varepsilon},$$

$$Q_\mu = \frac{X_\mu}{U_\mu}, \quad U_\mu = \prod_{\substack{\nu=1 \\ \nu \neq \mu}}^n (x_\mu^2 - x_\nu^2), \quad X_\mu = X_\mu(x_\mu),$$

$$A_\mu^{(k)} = \sum_{\substack{1 \leq \nu_1 < \dots < \nu_k \leq n \\ \nu_i \neq \mu}} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A^{(k)} = \sum_{1 \leq \nu_1 < \dots < \nu_k \leq n} x_{\nu_1}^2 \cdots x_{\nu_k}^2, \quad A_\mu^{(0)} = A^{(0)} = 1,$$

$$S = \frac{c}{A^{(n)}}, \quad c = \text{const.}$$



From the viewpoint of hidden symmetries, it is convenient to use a string-frame metric  $g_s$  which is conformally related to a Einstein-frame metric  $g_E$  by

$$g_E = X^{-2} g_s .$$

Then it leads to the string-frame Lagrangian

$$\mathcal{L}_D = X^{-(D-2)} \left\{ * R_s + \frac{1}{2} * d\varphi \wedge d\varphi - * F_{(2)} \wedge F_{(2)} - \frac{1}{2} * H_{(3)} \wedge H_{(3)} \right\} .$$

In string frame the metric  $g_s$  is written as

$$g_s = \sum_{\mu=1}^n (e^\mu e^\mu + e^{\hat{\mu}} e^{\hat{\mu}}) + \varepsilon e^0 e^0 ,$$

where the vielbeins for Chow's solution are

$$e^\mu = \frac{dx_\mu}{\sqrt{Q_\mu}} , \quad e^{\hat{\mu}} = \sqrt{Q_\mu} \left( \mathcal{A}_\mu - \sum_{\nu=1}^n \frac{2N_\nu s^2}{HU_\nu} \mathcal{A}_\nu \right) , \quad e^0 = \sqrt{S} \left( \mathcal{A} - \sum_{\nu=1}^n \frac{2N_\nu s^2}{HU_\nu} \mathcal{A}_\nu \right) .$$

As we find soon, there are  $n + \varepsilon$  Killing vectors given by  $\partial/\partial\psi_k$ ,  $k = 0, \dots, n - 1 + \varepsilon$ . In addition, it is known that there are  $n - 1$  rank-2 Killing tensors  $K^{(j)}$  given by

$$K^{(j)} = \sum_{\mu=1}^n A_\mu^{(j)} (e^\mu e^\mu + e^{\hat{\mu}} e^{\hat{\mu}}) + \varepsilon A^{(j)} e^0 e^0 ,$$

where  $j = 1, \dots, n - 1$ . Consequently, there are in Einstein frame  $n - 1$  rank-2 conformal Killing tensors  $Q^{(j)}$  given by

$$Q^{(j)} = H^{2/(D-2)} K^{(j)} .$$

# Generalized Closed Conformal Killing-Yano Tensor

Kubizňák, Kunduri and Yasui (2008)

**Def.** Let  $h$  be a  $p$ -form and  $T$  be a 3-form. If a pair of  $(h, T)$  satisfies the equations

$$\nabla_X^T h = -\frac{1}{D-p+1} X^b \wedge \delta^T h \quad \text{and} \quad d^T h = 0$$

for  $\forall X \in TM$ , then we call  $h$  **rank- $p$  generalized closed conformal Killing-Yano (GCCKY) tensor with 3-form  $T$** .

$\nabla$  : Levi-Civita connection,  $\wedge$  : wedge product,  $d$  : exterior derivative,  
 $\delta$  : coderivative operator ( $= *d*$ ),  $\lrcorner$  : inner product

$$\nabla_X^T h := \nabla_X h - \frac{1}{2} \sum_a (X \lrcorner e_a \lrcorner T) \wedge (e_a \lrcorner h) ,$$
$$d^T h := \sum_a e^a \wedge \nabla_{e_a}^T h , \quad \delta^T h := - \sum_a e_a \lrcorner \nabla_{e_a}^T h .$$

**Prop.** Let  $(M, g)$  be a  $D$ -dimensional spacetime. If  $(M, g)$  admits a rank-2 non-degenerate GCCKY tensor  $h$  with a 3-form  $T$  then there exist  $n - 1$  rank-2 Killing tensors  $K^{(j)}$  ( $j = 1, \dots, n - 1$ ).

$$h = \sum_{\mu=1}^n x_{\mu} e^{\mu} \wedge e^{\hat{\mu}}, \quad K^{(j)} = \sum_{\mu=1}^n A_{\mu}^{(j)} (e^{\mu} e^{\mu} + e^{\hat{\mu}} e^{\hat{\mu}}) + \varepsilon A^{(j)} e^0 e^0$$

$\{e^a\}$  : orthonormal basis

### Difference 1

With  $T = 0$  all commutators of Killing tensors vanish automatically, but with  $T \neq 0$  it doesn't occur.

### Difference 2

With  $T = 0$  rank-2 CCKY tensor leads to  $n + \varepsilon$  Killing vectors, but it doesn't occur with  $T \neq 0$ .

For geodesic integrability we need **some additional condition for  $T$** .

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$$S = \frac{c}{A^{(n)}}, \quad c = \text{const.}$$

For Chow's solution **in string frame**, we find a rank-2 GCCKY tensor

$$h = \sum_{\mu=1}^n x_{\mu} e^{\mu} \wedge e^{\hat{\mu}}$$

with a 3-form

$$h = \sum_{\rho=1}^n \sum_{\substack{\mu=1 \\ \mu \neq \rho}}^n \sqrt{Q_{\mu}} (\partial_{\rho} \ln H) e^{\rho} \wedge e^{\hat{\mu}} \wedge e^{\hat{\rho}} \\ - \varepsilon \sum_{\rho=1}^n \sqrt{S} (\partial_{\rho} \ln H) e^{\rho} \wedge e^{\hat{\rho}} \wedge e^0 + \varepsilon \sum_{\rho=1}^n \frac{f}{x_{\rho}} e^{\rho} \wedge e^{\hat{\rho}} \wedge e^0 ,$$

where  $f$  is an arbitrary function.

When  $f = 0$ , we can write the 3-form  $T$  as

$$T = k X^{D-6} H_{(3)} ,$$

where  $H_{(3)} = dB_{(2)} - A_{(1)} \wedge dA_{(1)}$  and  $k$  is some constant.

Thus **in string frame** there are  $n - 1$  rank-2 Killing tensors  $K^{(j)}$  given by

$$K^{(j)} = \sum_{\mu=1}^n A_{\mu}^{(j)} (e^{\mu} e^{\mu} + e^{\hat{\mu}} e^{\hat{\mu}}) + \varepsilon A^{(j)} e^0 e^0 ,$$

where  $j = 1, \dots, n - 1$ .

One can check that the torsion  $T$  satisfies a condition on which Killing tensors are mutually commuting.

Consequently, there are **in Einstein-frame**  $n - 1$  rank-2 conformal Killing tensors  $Q^{(j)}$  given by

$$Q^{(j)} = H^{2/(D-2)} K^{(j)} .$$

# Summary

- We have introduced the notion of (G)CCKY tensor and showed the relation to geodesic integrability.
- By imposing a rank-2 non-degenerate CCKY tensor we have constructed a metric ansatz which has geodesic integrability and examined solutions to (vacuum) Einstein equation.
- We have considered the charged Kerr-NUT spacetime given by Chow's solution, which includes ...

**Kerr-Sen black hole in 4 dimension,**

**charged rotating black hole with  $\delta_1 = \delta_2$  and  $\delta_3 = 0$  in 5-dim.  
 $U(1)^3$  ungauged supergravity, etc.**

- We have understood that the Killing tensors for the charged Kerr-NUT spacetime (in string frame) come from a rank-2 GCCKY tensor.



## **Discussion** - small questions -

- Properties of charged Kerr-NUT spacetime
  - relation between GCCKY tensor and Killing vectors?**
  - separability of Klein-Gordon equation, Dirac equation, etc?**
- How about other known solutions?
- General properties of GCCKY tensor
  - What's the condition that Killing vectors can be constructed from a GCCKY tensor?**
  - Are symmetry operators which commute with Laplacian, Dirac operator, etc constructed from it?**

## Discussion - large questions -

- Can we construct new solution?

e.g. vacuum black hole solutions, [Houri, Oota and Yasui \(2007\)](#)  
black hole solution in 5-dim. minimal gauged supergravity  
[Ahmedov and Aliev \(2009\)](#)

- What's the physical meaning?
- Why many known black hole solutions have such a symmetry?

e.g. black ring solution doesn't admit CCKY tensor.

It seems to me that these questions are deeply related each other...