# Wobbling kinks in a coupled two-component $\phi^4$ field theory

A. Alonso-Izquierdo<sup>1</sup>, D. Miguélez-Caballero<sup>2</sup>, L.M. Nieto<sup>2</sup>

<sup>1</sup> Department of Applied Mathematics - University of Salamanca
<sup>2</sup> Department of Theoretical, Optical and Atomic Physics - University of Valladolid david.miguelez@uva.es

September 13<sup>th</sup>, 2023



Universidad de Valladolid



**ĐSALAMANCA** 

### Outline

- Among the range of topological solitons, kinks are the simplest and arise in a wide variety of scalar field theories.
- We will study one of the kinks that arise in a two-component  $\phi^4$  scalar field theory coupled by means of a parameter  $\kappa$ .
- We will unravel the behaviour of the kink solution when one of its shape modes is initially triggered. We will use a perturbative approach and then we will compare the results obtained with numerical simulations.

- Among the range of topological solitons, kinks are the simplest and arise in a wide variety of scalar field theories.
- We will study one of the kinks that arise in a two-component  $\phi^4$  scalar field theory coupled by means of a parameter  $\kappa$  .
- We will unravel the behaviour of the kink solution when one of its shape modes is initially triggered. We will use a perturbative approach and then we will compare the results obtained with numerical simulations.

- Among the range of topological solitons, kinks are the simplest and arise in a wide variety of scalar field theories.
- We will study one of the kinks that arise in a two-component  $\phi^4$  scalar field theory coupled by means of a parameter  $\kappa$  .
- We will unravel the behaviour of the kink solution when one of its shape modes is initially triggered. We will use a perturbative approach and then we will compare the results obtained with numerical simulations.

### 2.Kinks in the two-component $\phi^4$ model

ullet The model we are going to study is given by the Lagrangian density  $^1$ 

$$\mathcal{L} = rac{1}{2} \partial_\mu \phi \partial^\mu \phi + rac{1}{2} \partial_\mu \psi \partial^\mu \psi - U(\phi, \psi).$$

where

$$U(\phi,\psi) = \frac{1}{2}(\phi^2 - 1)^2 + \frac{1}{2}(\psi^2 - 1)^2 + \kappa \phi^2 \psi^2 - \frac{1}{2}.$$

• The associated field equations are

$$\begin{aligned} \partial_{tt}\phi - \partial_{xx}\phi + 2\phi(\phi^2 - 1 + \kappa\psi^2) &= 0, \\ \partial_{tt}\psi - \partial_{xx}\psi + 2\psi(\psi^2 - 1 + \kappa\phi^2) &= 0. \end{aligned}$$

• For the potential  $U(\phi,\psi)$  the vacua structure depends on the value of  $\kappa$ :

$$\begin{split} \mathcal{M}_{\kappa < 1} &= \left\{ \frac{1}{\sqrt{1+\kappa}} \begin{pmatrix} (-1)^a \\ (-1)^b \end{pmatrix}, \quad a, b = 0, 1 \right\}, \\ \mathcal{M}_{\kappa > 1} &= \left\{ \begin{pmatrix} (-1)^a \\ 0 \end{pmatrix} \begin{pmatrix} 0 \\ (-1)^b \end{pmatrix}, \quad a, b = 0, 1 \right\}. \end{split}$$

<sup>1</sup>A. Halavanau, T. Romanczukiewicz and Y.M. Shnir, *Resonance structures in coupled two-component*  $\phi^4$  *model*, Phys. Rev. D **86**, 085027 (2012).

#### 2.Kinks in the two-component $\phi^4$ model

ullet The model we are going to study is given by the Lagrangian density  $^1$ 

$$\mathcal{L} = rac{1}{2} \partial_\mu \phi \partial^\mu \phi + rac{1}{2} \partial_\mu \psi \partial^\mu \psi - U(\phi, \psi).$$

where

$$U(\phi,\psi) = \frac{1}{2}(\phi^2 - 1)^2 + \frac{1}{2}(\psi^2 - 1)^2 + \frac{\kappa\phi^2\psi^2}{2} - \frac{1}{2}.$$

• The associated field equations are

$$\begin{aligned} \partial_{tt}\phi - \partial_{xx}\phi + 2\phi(\phi^2 - 1 + \kappa\psi^2) &= 0, \\ \partial_{tt}\psi - \partial_{xx}\psi + 2\psi(\psi^2 - 1 + \kappa\phi^2) &= 0. \end{aligned}$$

• For the potential  $U(\phi,\psi)$  the vacua structure depends on the value of  $\kappa$ :

$$\begin{split} \mathcal{M}_{\kappa < 1} &= \left\{ \frac{1}{\sqrt{1+\kappa}} \begin{pmatrix} (-1)^a \\ (-1)^b \end{pmatrix}, \quad a, b = 0, 1 \right\}, \\ \mathcal{M}_{\kappa > 1} &= \left\{ \begin{pmatrix} (-1)^a \\ 0 \end{pmatrix} \begin{pmatrix} 0 \\ (-1)^b \end{pmatrix}, \quad a, b = 0, 1 \right\}. \end{split}$$

<sup>1</sup>A. Halavanau, T. Romanczukiewicz and Y.M. Shnir, *Resonance structures in coupled two-component*  $\phi^4$  *model*, Phys. Rev. D **86**, 085027 (2012).

### 2.Kinks in the two-component $\phi^4$ model

ullet The model we are going to study is given by the Lagrangian density  $^1$ 

$$\mathcal{L} = rac{1}{2} \partial_\mu \phi \partial^\mu \phi + rac{1}{2} \partial_\mu \psi \partial^\mu \psi - U(\phi, \psi).$$

where

$$U(\phi,\psi) = \frac{1}{2}(\phi^2 - 1)^2 + \frac{1}{2}(\psi^2 - 1)^2 + \frac{\kappa\phi^2\psi^2}{2} - \frac{1}{2}.$$

• The associated field equations are

$$\begin{aligned} \partial_{tt}\phi - \partial_{xx}\phi + 2\phi(\phi^2 - 1 + \kappa\psi^2) &= 0, \\ \partial_{tt}\psi - \partial_{xx}\psi + 2\psi(\psi^2 - 1 + \kappa\phi^2) &= 0. \end{aligned}$$

• For the potential  $U(\phi, \psi)$  the vacua structure depends on the value of  $\kappa$ :

$$\begin{split} \mathcal{M}_{\kappa < 1} &= \left\{ \frac{1}{\sqrt{1+\kappa}} \begin{pmatrix} (-1)^{\mathfrak{a}} \\ (-1)^{b} \end{pmatrix}, \quad \mathfrak{a}, b = 0, 1 \right\}, \\ \mathcal{M}_{\kappa > 1} &= \left\{ \begin{pmatrix} (-1)^{\mathfrak{a}} \\ 0 \end{pmatrix} \begin{pmatrix} 0 \\ (-1)^{b} \end{pmatrix}, \quad \mathfrak{a}, b = 0, 1 \right\}. \end{split}$$

<sup>1</sup>A. Halavanau, T. Romanczukiewicz and Y.M. Shnir, *Resonance structures in coupled two-component*  $\phi^4$  *model*, Phys. Rev. D **86**, 085027 (2012).

# 2.kinks in the two-component $\phi^4$ model

• <u>*\kappa* < 1</u>

Kink solutions take the form:

$$\mathcal{K}^{(a,b)}(x) = \frac{\tanh x}{\sqrt{1+\kappa}} \begin{pmatrix} (-1)^a \\ (-1)^b \end{pmatrix}, \quad a, b = 0, 1.$$

Shown over the potential:



# 2.kinks in the two-component $\phi^4$ model

<u>κ > 1</u>

Kink solutions take the form:

$$\mathcal{K}_{1}^{(\pm)}(x) = \begin{pmatrix} \pm \tanh x \\ 0 \end{pmatrix}, \quad \mathcal{K}_{2}^{(\pm)}(x) = \begin{pmatrix} 0 \\ \pm \tanh x \end{pmatrix}$$

Shown over the potential:



• When the stability of a kink solution is studied we have to propose a solution with the structure

$$\widetilde{K}(x,t;\omega,a) = K_1^{(\pm)}(x) + a e^{i\omega t} F_{\omega}(x).$$

where  $K_1^{(\pm)}(x)$  is the kink solution and *a* is a small real parameter. Plugging this solution would lead to the problem

$$\mathcal{H} F_{\omega}(x) = \omega^2 F_{\omega}(x)$$

where

$$\mathcal{H} = \left( \begin{array}{cc} -\frac{d^2}{dx^2} + 4 - 6 \operatorname{sech}^2 x & 0 \\ 0 & -\frac{d^2}{dx^2} + 2\kappa \tanh^2 x - 2 \end{array} \right)$$

• When the stability of a kink solution is studied we have to propose a solution with the structure

$$\widetilde{K}(x,t;\omega,a) = K_1^{(\pm)}(x) + a e^{i\omega t} F_{\omega}(x).$$

where  $K_1^{(\pm)}(x)$  is the kink solution and *a* is a small real parameter.

• Plugging this solution would lead to the problem

$$\mathcal{H} F_{\omega}(x) = \omega^2 F_{\omega}(x)$$

where

$$\mathcal{H} = \begin{pmatrix} -\frac{d^2}{dx^2} + 4 - 6 \operatorname{sech}^2 x & 0 \\ 0 & -\frac{d^2}{dx^2} + 2\kappa \tanh^2 x - 2 \end{pmatrix}$$

#### Longitudinal eigenmodes

Eigenfrequency $(\omega)$	Eigenfunction
0	$(\overline{\eta}_0(x),0)^t=(\mathrm{sech}^2x,0)^t$
$\overline{\omega} = \sqrt{3}$	$(\overline{\eta}_D(x),0)^t = (\operatorname{sech} x \operatorname{tanh} x,0)^t$
$\overline{\omega}^c_{\overline{q}} = \sqrt{4+\overline{q}^2}$	$\left(\overline{\eta}_{\overline{q}}(x),0\right)^{t}=\left(e^{i\overline{q}x}\left[-1-\overline{q}^{2}+3\tanh^{2}x-3i\overline{q}\tanh x\right],0\right)^{t}$

#### Orthogonal eigenmodes

where  $\overline{q}$  and  $\widehat{q}$  are real quantities,  $\rho = \sqrt{2\kappa + \frac{1}{4}}$  and *n* is a natural number whose maximum value is given by the relation  $\kappa > \frac{n_{\max}(n_{\max}+1)}{2}$ . For  $\kappa < 3$ ,  $\widehat{\omega}_{D,0}^2 < 0$ , which means that the kink is unstable in this regime.

#### Longitudinal eigenmodes

Eigenfrequency $(\omega)$	Eigenfunction
0	$(\overline{\eta}_0(x),0)^t=(\mathrm{sech}^2x,0)^t$
$\overline{\omega} = \sqrt{3}$	$(\overline{\eta}_D(x),0)^t = (\operatorname{sech} x \operatorname{tanh} x,0)^t$
$\overline{\omega}^{ extsf{c}}_{\overline{q}} = \sqrt{4+\overline{q}^2}$	$(\overline{\eta}_{\overline{q}}(x),0)^{t} = \left(e^{i\overline{q}x}\left[-1-\overline{q}^{2}+3 \tanh^{2}x-3i\overline{q} \tanh x\right],0\right)^{t}$

#### **Orthogonal eigenmodes**

Eigenfrequency $(\omega)$	Eigenfunction
$\hat{\omega}_{D,n} = \sqrt{(2n+1)\rho - n^2 - n - \frac{5}{2}}$	$(0, \widehat{\eta}_D(x))^t = (0, (\operatorname{sech} x)^{\rho - n - \frac{1}{2}}  _2F_1(-n, 2\rho - n,  \rho - n + 1/2,  \frac{(1 - \tanh x)}{2}))^t$
$\widehat{\omega}_{\widehat{q}}^{c} = \sqrt{\widehat{q}^2 + 2\kappa - 2}$	$\left(0 \ , \ \ \widehat{\eta}_{\widehat{q}}(x)\right)^{t} = \left(0, \ \ _{2}F_{1}(\frac{1}{2} - \rho, \frac{1}{2} + \rho, -i\widehat{q} + 1, \frac{(1-\tanh x)}{2}) \ e^{i\widehat{q}x}\right)^{t}$

where  $\overline{q}$  and  $\widehat{q}$  are real quantities,  $\rho = \sqrt{2\kappa + \frac{1}{4}}$  and *n* is a natural number whose maximum value is given by the relation  $\kappa > \frac{n_{\max}(n_{\max}+1)}{2}$ . For  $\kappa < 3$ ,  $\widehat{\omega}_{D,0}^2 < 0$ , which means that the kink is unstable in this regime.

- We are interested in studying the evolution of the system when one orthogonal mode is initially activated.
- We are going to assume a general expansion of the fields of the form <sup>2</sup>

$$\begin{split} \phi(\mathbf{x},t) &= \phi_{\mathcal{K}}(\mathbf{x}) + \overline{a}(t) \, \overline{\eta}_{D}(\mathbf{x}) + \overline{\eta}(\mathbf{x},t), \\ \psi(\mathbf{x},t) &= \sum_{p} \widehat{a}_{p}(t) \, \widehat{\eta}_{D,p}(\mathbf{x}) + \widehat{\eta}(\mathbf{x},t), \end{split}$$

where  $\phi_K(x) = \tanh x$ . Plugging this assumption into the field equations and neglecting the smallest terms we find

$$\begin{aligned} \left(\overline{a}_{tt} + \overline{\omega}^2 \,\overline{a}\right) \overline{\eta}_D + \overline{\eta}_{tt} - \overline{\eta}_{xx} - 2\overline{\eta} + 6\,\overline{\eta}\,\phi_K^2 + 6\overline{a}^2\,\overline{\eta}_D^2\,\phi_K + 2\kappa\,\phi_K\,(\sum_p \widehat{a}_p\widehat{\eta}_{D,p})^2 \approx 0, \\ \sum_p \left(\left(\widehat{a}_{tt}\right)_p + \widehat{\omega}_p^2\,\widehat{a}_p\right)\widehat{\eta}_{D,p} + \widehat{\eta}_{tt} - \widehat{\eta}_{xx} - 2\widehat{\eta} + 2\kappa\,\widehat{\eta}\,\phi_K^2 + 4\kappa\,\overline{a}\,\overline{\eta}_D\,\phi_K\sum_p \widehat{a}_p\,\widehat{\eta}_{D,p} \approx 0. \end{aligned}$$

<sup>&</sup>lt;sup>2</sup>N.S. Manton and H. Merabet, *Kinks-gradient flow and dynamics*, Nonlinearity 10, 3 (1997).

9/1

- We are interested in studying the evolution of the system when one orthogonal mode is initially activated.
- We are going to assume a general expansion of the fields of the form <sup>2</sup>

$$\begin{split} \phi(x,t) &= \phi_{\mathcal{K}}(x) + \overline{a}(t) \, \overline{\eta}_D(x) + \overline{\eta}(x,t), \\ \psi(x,t) &= \sum_p \widehat{a}_p(t) \, \widehat{\eta}_{D,p}(x) + \widehat{\eta}(x,t), \end{split}$$

where  $\phi_K(x) = \tanh x$ . Plugging this assumption into the field equations and neglecting the smallest terms we find

$$\begin{aligned} &(\overline{a}_{tt} + \overline{\omega}^2 \,\overline{a}) \,\overline{\eta}_D + \overline{\eta}_{tt} - \overline{\eta}_{xx} - 2\overline{\eta} + 6 \,\overline{\eta} \,\phi_K^2 + 6\overline{a}^2 \,\overline{\eta}_D^2 \,\phi_K + 2\kappa \,\phi_K \,(\sum_p \widehat{a}_p \widehat{\eta}_{D,p})^2 \approx 0, \\ &\sum_p ((\widehat{a}_{tt})_p + \widehat{\omega}_p^2 \,\widehat{a}_p) \,\widehat{\eta}_{D,p} + \widehat{\eta}_{tt} - \widehat{\eta}_{xx} - 2\widehat{\eta} + 2\kappa \,\widehat{\eta} \,\phi_K^2 + 4\kappa \,\overline{a} \,\overline{\eta}_D \,\phi_K \sum_p \widehat{a}_p \,\widehat{\eta}_{D,p} \approx 0. \end{aligned}$$

<sup>2</sup>N.S. Manton and H. Merabet, *Kinks-gradient flow and dynamics*, Nonlinearity **10**, 3 (1997).

If the previous equations are projected onto  $\overline{\eta}$  and  $\widehat{\eta}_{D,m},$  then:

$$(\overline{a}_{tt} + \overline{\omega}^2 \,\overline{a}) \,\overline{C}_D^2 + 6 \,\overline{a}^2 \,\overline{V} + \sum_{p,r} \widehat{a}_p \,\widehat{a}_r \,\widehat{B}_{pr} = 0,$$
$$((\widehat{a}_{tt})_m + \widehat{\omega}_m^2 \,\widehat{a}_m) \,\widehat{C}_{D,m}^2 + 2 \sum_p \overline{a} \,\widehat{a}_p \,\widehat{B}_{pm} = 0.$$

As it is assumed that we initially trigger the *j*-th shape mode, then

 $\widehat{a}_j(t) \approx a_0 \sin(\widehat{\omega}_j t).$ 

If we now plug this formula into the first differential equation with the initial conditions

$$\overline{a}_t(0) = \overline{a}(0) = 0, \quad \widehat{a}_m(0) = \widehat{a}_m(0)_t = 0 \quad ext{with} \quad m 
eq j,$$

we have that

$$\overline{a}(t) pprox rac{a_0^2 \widehat{B}_{jj} \left(4 \widehat{\omega}_j^2 - \overline{\omega}^2 + \overline{\omega}^2 \cos(2 \widehat{\omega}_j t) - 4 \widehat{\omega}_j^2 \cos(\overline{\omega} t)
ight)}{2 \overline{C}_D^2 \overline{\omega}^2 (\overline{\omega}^2 - 4 \widehat{\omega}_j^2)}$$

If the previous equations are projected onto  $\overline{\eta}$  and  $\widehat{\eta}_{D,m},$  then:

$$(\overline{a}_{tt} + \overline{\omega}^2 \,\overline{a}) \,\overline{C}_D^2 + 6 \,\overline{a}^2 \,\overline{V} + \sum_{p,r} \widehat{a}_p \,\widehat{a}_r \,\widehat{B}_{pr} = 0,$$
$$((\widehat{a}_{tt})_m + \widehat{\omega}_m^2 \,\widehat{a}_m) \,\widehat{C}_{D,m}^2 + 2 \sum_p \overline{a} \,\widehat{a}_p \,\widehat{B}_{pm} = 0.$$

As it is assumed that we initially trigger the *j*-th shape mode, then  $\widehat{a_j}(t) \approx a_0 \sin(\widehat{\omega}_j t).$ 

If we now plug this formula into the first differential equation with the initial conditions

$$\overline{a}_t(0) = \overline{a}(0) = 0, \quad \widehat{a}_m(0) = \widehat{a}_m(0)_t = 0 \quad ext{with} \quad m \neq j,$$

we have that

$$\overline{a}(t) \approx \frac{a_0^2 \widehat{B}_{jj} \left(4 \widehat{\omega}_j^2 - \overline{\omega}^2 + \overline{\omega}^2 \cos(2 \widehat{\omega}_j t) - 4 \widehat{\omega}_j^2 \cos(\overline{\omega} t)\right)}{2 \overline{C}_D^2 \overline{\omega}^2 (\overline{\omega}^2 - 4 \widehat{\omega}_j^2)}$$

If the previous equations are projected onto  $\overline{\eta}$  and  $\widehat{\eta}_{D,m},$  then:

$$(\overline{a}_{tt} + \overline{\omega}^2 \,\overline{a}) \,\overline{C}_D^2 + 6 \,\overline{a}^2 \,\overline{V} + \sum_{p,r} \widehat{a}_p \,\widehat{a}_r \,\widehat{B}_{pr} = 0,$$
$$((\widehat{a}_{tt})_m + \widehat{\omega}_m^2 \,\widehat{a}_m) \,\widehat{C}_{D,m}^2 + 2 \sum_p \overline{a} \,\widehat{a}_p \,\widehat{B}_{pm} = 0.$$

As it is assumed that we initially trigger the *j*-th shape mode, then

 $\widehat{a}_j(t) \approx a_0 \sin(\widehat{\omega}_j t).$ 

If we now plug this formula into the first differential equation with the initial conditions

$$\overline{a}_t(0) = \overline{a}(0) = 0, \quad \widehat{a}_m(0) = \widehat{a}_m(0)_t = 0 \quad \text{with} \quad m \neq j,$$

we have that

$$\overline{a}(t)pprox rac{a_0^2\widehat{B}_{jj}\left(4\widehat{\omega}_j^2-\overline{\omega}^2+\overline{\omega}^2\cos(2\widehat{\omega}_jt)-4\widehat{\omega}_j^2\cos(\overline{\omega}t)
ight)}{2\overline{\mathcal{C}}_D^2\overline{\omega}^2(\overline{\omega}^2-4\widehat{\omega}_j^2)}.$$

Using the previous relations in the original truncated expansion the next differential equations are found:

$$\begin{aligned} &-\overline{\eta}^{\,\prime\prime}(x) + \left(6\phi_K^2 - 2 - 4\widehat{\omega}_j^2\right)\overline{\eta}(x) &= f(x),\\ &-\widehat{\eta}^{\,\prime\prime}(x) + \left(2\kappa\,\phi_K^2 - 2 - \omega_\ell^2\right)\widehat{\eta}(x) &= g_\ell(x), \end{aligned}$$

where  $\ell = 1, 2$  and  $\omega_1 = 3\widehat{\omega}_j$ ,  $\omega_2 = \widehat{\omega}_j + \overline{\omega}$ .

These equations describe the radiation emitted on the longitudinal channel at frequency  $2\hat{\omega}_j$  and on the orthogonal one at frequencies  $3\hat{\omega}_j$  and  $\hat{\omega}_j + \overline{\omega}$ .

Using the previous relations in the original truncated expansion the next differential equations are found:

$$\begin{aligned} &-\overline{\eta}^{\,\prime\prime}(x) + \left(6\phi_K^2 - 2 - 4\widehat{\omega}_j^2\right)\overline{\eta}(x) &= f(x),\\ &-\widehat{\eta}^{\,\prime\prime}(x) + \left(2\kappa\,\phi_K^2 - 2 - \omega_\ell^2\right)\widehat{\eta}(x) &= g_\ell(x), \end{aligned}$$

where  $\ell = 1, 2$  and  $\omega_1 = 3\widehat{\omega}_j$ ,  $\omega_2 = \widehat{\omega}_j + \overline{\omega}$ .

These equations describe the radiation emitted on the longitudinal channel at frequency  $2\hat{\omega}_i$  and on the orthogonal one at frequencies  $3\hat{\omega}_i$  and  $\hat{\omega}_i + \overline{\omega}$ .

Using the **variation of parameters** method the assymptotic behaviour of the radiation terms can be computed:

$$ar{\eta}_{2\widehat{\omega}_{j}} \quad \xrightarrow{x o \infty} \quad rac{i\left(\int_{-\infty}^{\infty} \overline{\eta}_{\overline{q}}(y)f(y)dy
ight)}{2(\overline{q}+i)(\overline{q}+2i)}e^{-i\overline{q}x}, \ \widehat{\eta}_{\widehat{\omega}_{\ell}} \quad \xrightarrow{x o \infty} \quad rac{\left(\int_{-\infty}^{\infty} \widehat{\eta}_{\widehat{q}_{\ell}}(y)g_{\ell}(y)dy
ight)}{2i\widehat{q}_{\ell}}e^{-i\widehat{q}_{\ell}x}.$$

where

$$\overline{\mathbf{q}} = 2\sqrt{\widehat{\omega}_j^2 - 1},$$

$$\widehat{q}_1 = \sqrt{9\widehat{\omega}_j^2 + 2 - 2\kappa},$$

$$\widehat{q}_2 = \sqrt{(\widehat{\omega}_j + \overline{\omega})^2 + 2 - 2\kappa}$$



#### 4. Numerical results: Radiation amplitudes



#### 14/1

#### 4. Numerical results: Radiation amplitudes

#### **Orthogonal Radiation Amplitudes**



• We have assumed that the amplitude of the triggered shape mode remains constant. This assumption works fine when  $\hat{\eta}_{D,1}$  and  $\hat{\eta}_{D,2}$  are initially activated.



• But for  $\hat{\eta}_{D,0}$  a large decrease in the wobbling amplitude can be appreciated.



• A decay law for  $a_0$  can be computed in order. First, if we trigger  $\hat{\eta}_{D,0}$  the total radiated energy flux is given by

$$\langle P \rangle = rac{dE}{dt} = -(a_0^2 \, \overline{A}'_{2\widehat{\omega}_0})^2 \, (2 \, \widehat{\omega}_0) \, \overline{q},$$

where  $\overline{A}_{2\widehat{\omega}_0} = a_0^2 \ \overline{A}'_{2\widehat{\omega}_0}$ .

 On the other hand, the wobbling amplitude behaves as an harmonic oscillator on each point of the space so,

$$\mathcal{E} = \frac{1}{2}\widehat{\omega}_0^2 a_0^2 \widehat{\eta}_0^2 \rightarrow E = \int_{-\infty}^{\infty} \mathcal{E} \ d\mathbf{x} = \frac{1}{2} \ \widehat{\omega}_0^2 \ a_0^2 \ \widehat{C}_{D,0}^2.$$

• This leads to the following differential equation for  $a_0(t)$ 

$$\frac{1}{2}\,\widehat{\omega}_0^2\,\,\widehat{C}_{D,0}^2\,\,\frac{da_0^2(t)}{dt}\approx-2\widehat{\omega}_0\,\overline{A}_{2\widehat{\omega}_0}^{\prime 2}\,\overline{q}\,a_0^4(t),$$

whose solution is

$$a_0(t) pprox rac{a_0(0)}{\sqrt{1+t\left(rac{4\,\overline{q}\,a_0(0)^2\,\overline{A}_{2\widehat{\omega}_0}'^2}{\widehat{C}_{D,0}^2\,\widehat{\omega}_0}
ight)}}.$$

• A decay law for  $a_0$  can be computed in order. First, if we trigger  $\hat{\eta}_{D,0}$  the total radiated energy flux is given by

$$\langle P \rangle = \frac{dE}{dt} = -(a_0^2 \,\overline{A}'_{2\widehat{\omega}_0})^2 \,(2\,\widehat{\omega}_0)\,\overline{q},$$

where  $\overline{A}_{2\widehat{\omega}_0} = a_0^2 \ \overline{A}'_{2\widehat{\omega}_0}$ .

• On the other hand, the wobbling amplitude behaves as an harmonic oscillator on each point of the space so,

$$\mathcal{E} = \frac{1}{2}\widehat{\omega}_0^2 a_0^2 \widehat{\eta}_0^2 \rightarrow E = \int_{-\infty}^{\infty} \mathcal{E} \ dx = \frac{1}{2} \ \widehat{\omega}_0^2 \ a_0^2 \ \widehat{C}_{D,0}^2.$$

This leads to the following differential equation for a<sub>0</sub>(t)

$$\frac{1}{2}\,\widehat{\omega}_0^2\,\,\widehat{C}_{D,0}^2\,\,\frac{da_0^2(t)}{dt}\approx-2\widehat{\omega}_0\,\overline{A}_{2\widehat{\omega}_0}^{\prime 2}\,\overline{q}\,a_0^4(t),$$

whose solution is

$$a_0(t) pprox rac{a_0(0)}{\sqrt{1+t\left(rac{4\,\overline{q}\,a_0(0)^2\,\overline{A}_{2\widehat{\omega}_0}'^2}{\widehat{C}_{D,0}^2\,\widehat{\omega}_0}
ight)}}.$$

17/1

• A decay law for  $a_0$  can be computed in order. First, if we trigger  $\hat{\eta}_{D,0}$  the total radiated energy flux is given by

$$\langle P \rangle = \frac{dE}{dt} = -(a_0^2 \,\overline{A}'_{2\widehat{\omega}_0})^2 \,(2\,\widehat{\omega}_0)\,\overline{q},$$

where  $\overline{A}_{2\widehat{\omega}_0} = a_0^2 \ \overline{A}'_{2\widehat{\omega}_0}$ .

• On the other hand, the wobbling amplitude behaves as an harmonic oscillator on each point of the space so,

$$\mathcal{E} = \frac{1}{2}\widehat{\omega}_0^2 a_0^2 \widehat{\eta}_0^2 \rightarrow E = \int_{-\infty}^{\infty} \mathcal{E} \ dx = \frac{1}{2} \ \widehat{\omega}_0^2 \ a_0^2 \ \widehat{C}_{D,0}^2.$$

• This leads to the following differential equation for  $a_0(t)$ 

$$rac{1}{2}\,\widehat{\omega}_0^2\,\,\widehat{\mathcal{C}}_{D,0}^2\,\,rac{d \mathsf{a}_0^2(t)}{dt}pprox -2\widehat{\omega}_0\,\overline{\mathcal{A}}_{2\widehat{\omega}_0}^{\prime 2}\,\overline{q}\,\mathsf{a}_0^4(t),$$

whose solution is

$$a_0(t)pprox rac{a_0(0)}{\sqrt{1+t\left(rac{4\,\overline{q}\,a_0(0)^2\,\overline{A}_{2\widehat{\omega}_0}'}{\widehat{C}_{D,0}^2\,\widehat{\omega}_0}
ight)}}.$$

- Along this work, we have discussed the behaviour of an excited kink. In this process, radiation at three different frequencies has been found.
- Apart from this, the amplitudes found analytically predict the behaviour seen through numerical simulations.
- All the analytical methods used in this presentation can be extended in order to study other excited topological solitons such as Abelian-Higgs vortices.

- Along this work, we have discussed the behaviour of an excited kink. In this process, radiation at three different frequencies has been found.
- Apart from this, the amplitudes found analytically predict the behaviour seen through numerical simulations.
- All the analytical methods used in this presentation can be extended in order to study other excited topological solitons such as Abelian-Higgs vortices.

- Along this work, we have discussed the behaviour of an excited kink. In this process, radiation at three different frequencies has been found.
- Apart from this, the amplitudes found analytically predict the behaviour seen through numerical simulations.
- All the analytical methods used in this presentation can be extended in order to study other excited topological solitons such as Abelian-Higgs vortices.

# Thanks for your attention!!!